

Evolution of MHD Instabilities in Plasma Imploding Under Magnetic Field

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Abstract—Two-dimensional 3-ns-gated visible images, recorded at different times during the implosion of plasma under azimuthal magnetic field (*Z*-pinch), revealed ringlike instabilities followed by the development of axially and azimuthally nonuniform structures in the imploding plasma. Remarkably, the evolution in time of all structures was found to be highly repeatable in different shots, which should allow, through 3-D magnetohydrodynamics modeling, for systematically studying the development in time of these complex phenomena and correlating them with the initial plasma parameters. The data are also used to infer the time-dependent outer plasma radius and plasma radial velocities.

Index Terms—Plasma implosion, visible imaging, *Z*-pinch.

THE *Z*-PINCH implosions serve as intense sources of X-ray radiation that are of broad interest in the research toward inertial confinement fusion and the study of matter under high energy density [1]. In such systems, a cylindrical plasma column (typically formed out of an array of fine wires or a puff of gas) radially implodes under the azimuthal magnetic field resulting from axially driven high currents. In this process, acceleration-driven instabilities play a major role in the magnetohydrodynamics (MHD) of the plasma, including causing complex current paths in the plasma, magnetic-field penetration into certain parts of the plasma, and nonuniform disturbances in the implosion [2].

In this paper, a neon puff ($\sim 70 \mu\text{g}/\text{cm}$) was imploded in 500 ns until stagnation on axis under a current pulse that rises to 500 kA [3]. The puff consists of a 38-mm-diameter outer shell and an ~ 5 -mm-diameter on-axis jet. The distance between

the cylindrical nozzle (which serves as the cathode) and the cylindrical anode was 9.5 mm. Planar laser-induced fluorescence measurements on a similar system, but with an additional cylindrical nozzle, showed that the azimuthal uniformity of the initial gas density distribution is better than $\pm 15\%$ [4].

An *f*/16 optics with $7.5\times$ demagnification and coupled to an intensified CCD, viewing the plasma in the radial direction, provided a submillimeter spatial resolution across the 38-mm-diameter plasma shell, together with a sufficient signal. The filter used transmitted visible light above 6000 \AA ($T > 40\%$), thus mainly recording continuum radiation except for the earliest phase of the implosion where NeI–II line emission was significant too. The emission is thus nearly proportional to the square of the plasma density. The 3-ns-gated images were taken at different times along the implosion phase until the plasma stagnates on axis. Shown in Fig. 1 are 14 images of the imploding shell, most of them 50 ns apart, starting at $t = -500$ ns and ending at $t = 0$ (when the plasma stagnates on axis).

Surprisingly, it was found that the complex structures evolving in the plasma are repeatable in the different shots. Their repeatability even allows for tracking the growth of the structures at specific *z*-locations down to $t \cong -100$ ns.

The initial factors that govern the start of the growth of the nonuniform structures of the imploding shell are not known yet. One possibility is that they result from nonmonotonic variation of the initial gas density as a function of *r* and *z*, originating from the interaction between the gas jets as they propagate in the *z*-direction. The pictures also show the growth of the plasma portions that escape the implosion, pointing out the unresolved important question on the magnetic-field penetration into those portions. These pictures can serve as a unique tool for systematic examinations of 3-D MHD computations for the understanding of the initial conditions that give rise to the evolving structures and the time development of such structures.

These images are also used to plot the plasma visible-emitting radius as a function of time. It is seen that the radial motion of the plasma starts at $t = -250$ ns and ends at $t = 0$ with $r \sim 1$ mm. The radial velocities at the last 20 ns of the implosion are found to be $\sim 1.5 \times 10^7$ cm/s. This value is about 40% lower than the final implosion velocity obtained at $r < 0.5$ mm using X-ray imaging. This is expected since, in our experiments [3], [5], only $\cong 20\%$ of the very leading edge of the imploding plasma constitutes the *K*-emitting portion of the stagnating plasma.

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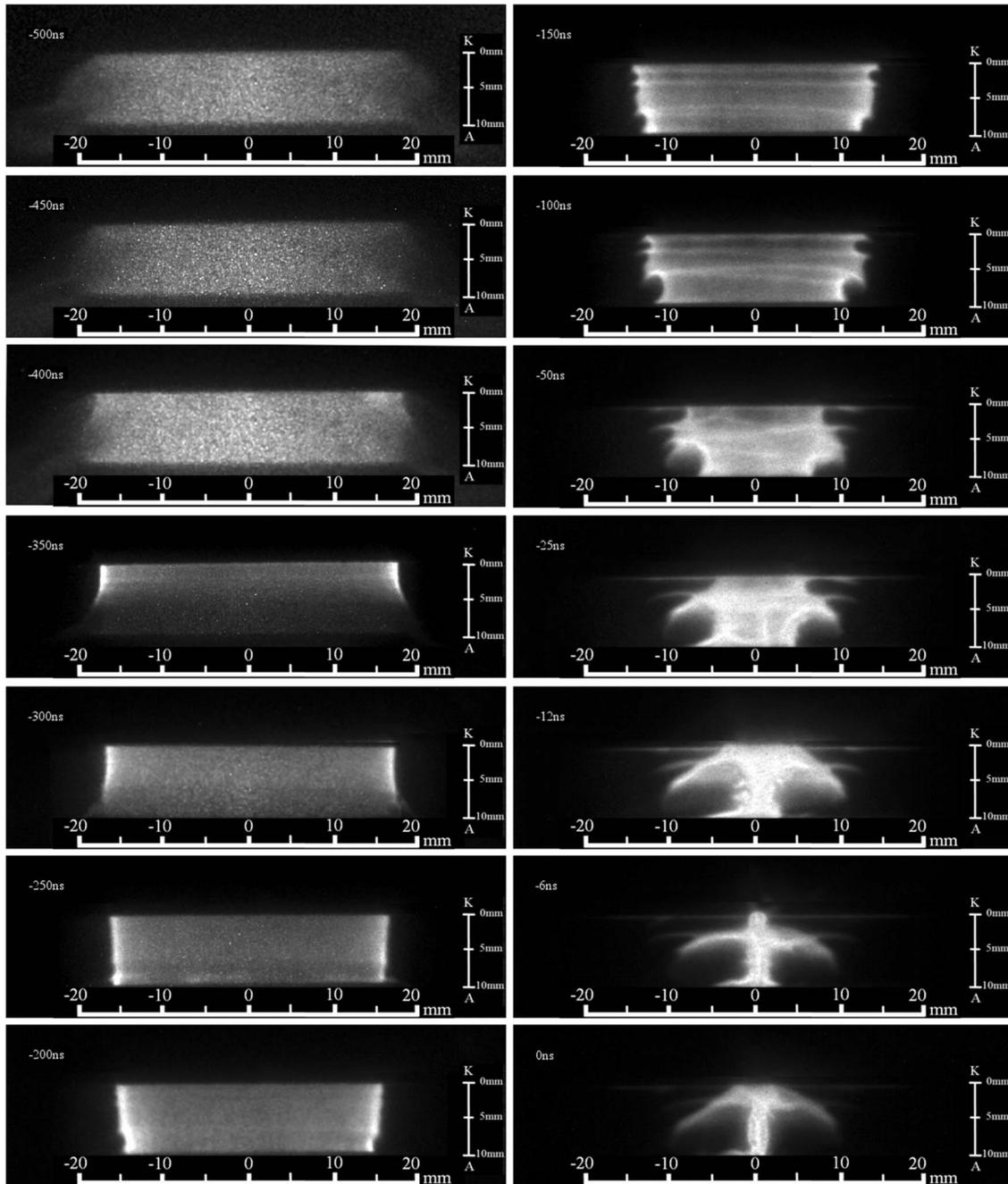


Fig. 1. Fourteen successive 2-D 3-ns-gated visible images of the imploding plasma under azimuthal magnetic field recorded at different times along the implosion phase. The times (where $t = 0$ is the pinch time) are indicated on the upper left corner of each image. The intensity of each image is normalized to the peak intensity of that frame.

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