



Communication Stark Broadening of Lyman-α in the Presence of a Strong Magnetic Field

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Abstract: Stark broadening of Lyman- α of a hydrogen-like atom in the presence of a strong magnetic field is analyzed. The shape of the central (π) component of the Lorentz–Zeeman triplet is expressed analytically, taking into account the plasma coupling and microfield dynamic effects. It is shown that in a sufficiently strong magnetic field, the broadening of this component, contrary to the broadening of the lateral (σ) ones, is independent of the magnetic field and, therefore, can be used for the plasma density diagnostics. Comparison with computer simulations at conditions typical for tokamak divertors and white dwarf atmospheres shows a very good agreement.

Keywords: line shapes; Zeeman effect; Stark effect

1. Introduction

Hydrogen atom is the simplest and best understood atomic system. It was the first real physical object to which the quantum mechanical description—first of the "old" Bohr's theory [1], then the modern quantum mechanics [2]—was applied and tested against.

The simplest transition in hydrogen or a hydrogen-like ion is Lyman- α ($n = 2 \rightarrow n = 1$, where n is the principal quantum number). Nevertheless, this transition represents a challenge for some models of plasma line broadening [3]. The reason is a strong influence of the ion dynamic effect, and in particular the directionality of the ion microfields [4] on the shape of the central Stark component of this line. Because of this effect, the Stark broadening of Lyman- α changes from the impact broadening [5] in the high-temperature/low-density regime to the so-called rotation broadening [6] in a low-temperature/high-density plasma, in a stark contrast to the majority of spectral lines that converge to the quasistatic lineshape [5].

In the presence of a sufficiently strong magnetic field, Lyman- α assumes a familiar Zeeman triplet pattern. However, its central (π , $\Delta M = 0$) and lateral (σ^{\pm} , $\Delta M = \pm 1$) components are broadened by the plasma Stark effect differently: Due to the degeneracy removal by the applied magnetic field, the upper states of the σ components, $|2p \pm 1\rangle$, are subject to the quadratic Stark effect, while that of the π component, $|2p0\rangle$, remains degenerate with $|2s0\rangle$ and, therefore, linearly depends on the electric field. As a result, the lateral components become significantly narrower than the central one [7].

2. Analytical Model

As an example, let us assume plasma conditions relevant to tokamak divertors. Specifically, the electron and ion temperature $T_e = T_i = 1 \text{ eV}$, the electron density $n_e = 10^{14} \text{ cm}^{-3}$, and the magnetic field *B* on the order of 1 T. A comparison of plasma-broadened Lyman- α shapes in the presence of the magnetic field is given in Figure 1 (see Section 3 for the details of these calculations). It is seen that for a non-zero magnetic field, the widths of the π and σ components of the Zeeman triplet are different, in agreement with the earlier findings [7]. Notably, as the magnetic-field strength increases, the Stark broadening of the central component remains nearly constant, whereas the lateral components become narrower.



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Figure 1. Stark-broadened Lyman- α shapes assuming $n_e = 10^{14} \text{ cm}^{-3}$, T = 1 eV, and a few values of the magnetic field as indicated in the legend.

To understand this result, consider Hamiltonian of the n = 2 manifold under the crossed \vec{F} and \vec{B} fields, e.g., see [8]:

$$V_{2} = \begin{pmatrix} 0 & 0 & f & 0 \\ 0 & -b\cos\theta & \frac{b}{\sqrt{2}}\sin\theta & 0 \\ f & \frac{b}{\sqrt{2}}\sin\theta & 0 & \frac{b}{\sqrt{2}}\sin\theta \\ 0 & 0 & \frac{b}{\sqrt{2}}\sin\theta & b\cos\theta \end{pmatrix},$$
 (1)

where for compactness $f := \frac{3}{Z}F$ and $b := \frac{\alpha}{2}B$ definitions are used, θ is the angle between \vec{F} and \vec{B} , and \vec{F} is assumed to lie along the quantization axis. Here, the spin degree of freedom is ignored (i.e., the magnetic-field perturbation is assumed to be much stronger than the fine structure). The atomic units ($\hbar = e = m_e = 1$) are used, α is the fine-structure constant, and Z is the charge of the nucleus (Z = 1 in the case of hydrogen).

The eigenvalues of V_2 are

$$(\Delta E)^2 = \frac{b^2 + f^2}{2} \pm \sqrt{\left(\frac{b^2 + f^2}{2}\right)^2 - f^2 b^2 \cos^2 \theta}.$$
 (2)

In the strong-*B* limit, i.e., $b \gg f$, this expression reduces to

$$\Delta E_{\pi} = \pm f \cos \theta \tag{3}$$

and

$$\Delta E_{\sigma} = \pm (b + \frac{f^2}{2b} \sin^2 \theta) \tag{4}$$

for the π and σ components, respectively. Thus, the Stark effect of the central component is linear (it is split into two sub-components), on the order of f, and independent of b, while that of the lateral components is quadratic, narrow ($\ll f$), and inversely proportional to b. The intensities of all four components, up to the $O((f/b)^2)$ terms, depend neither on the

The total area-normalized line shape $L(\omega)$ is a sum of the π and σ components which, after averaging over θ (but keeping *f* fixed) become, respectively,

$$L_{\pi}(\omega) = \begin{cases} \frac{1}{6f}, & |\omega| \le f\\ 0, & \text{elsewhere} \end{cases}$$
(5)

and

$$L_{\sigma}(\omega) = \begin{cases} \frac{b}{3f\sqrt{f^2 - 2b(|\omega| - b)}}, & b \le |\omega| < b + \frac{f^2}{2b} \\ 0, & \text{elsewhere.} \end{cases}$$
(6)

From Equation (5) it follows that $L_{\pi}(\omega)$ assumes a rectangular shape—a result recently obtained by [9]. Notably, this is the shape of any high-*n* H-like transition in the quasi-contiguous (QC) approximation [10]. Therefore, one can directly apply the QC-FFM approach [11] to obtain the π lineshape accounting for the microfield dynamics through the frequency-fluctuation model (FFM) [12,13]. Assuming the dynamic broadening of each of the π and σ^{\pm} components is independent [14], for a one-component plasma (OCP) one obtains

$$L_{\pi}(\bar{\nu};\bar{\omega}) = \frac{1}{\pi} \Re \frac{J(\bar{\nu};\bar{\omega})}{1 - \bar{\nu}J(\bar{\nu};\bar{\omega})} , \qquad (7)$$

where \Re stands for the real part and

$$J(\bar{\nu};\bar{\omega}) = \int_0^\infty d\tau \exp\left[-\phi(\tau) - i(\bar{\omega} - i\bar{\nu})\tau\right].$$
(8)

Here, the line shape is expressed as a function of the dimensionless reduced detuning $\bar{\omega} = \omega / \Delta_0$ and \bar{v} —a single parameter related to the typical frequency of the microfields in the radiator–perturber center-of-mass frame [15]

$$w_{\rm dyn} = \sqrt{\frac{kT_p}{m_p} + \frac{kT_r}{m_r}} \left(\frac{4\pi n_p}{3}\right)^{1/3} \tag{9}$$

via

$$\bar{\nu} = \frac{1}{2} \frac{w_{\rm dyn}}{\Delta_0} + \frac{1}{20} \left(\frac{w_{\rm dyn}}{\Delta_0}\right)^2,\tag{10}$$

where the second term is a semiempirical correction to recover the impact limit [5]. Z_p , n_p , m_p , and T_p are the charge, density, mass, and temperature of the OCP particles; m_r and T_r are the mass and temperature of the radiator. The normal detuning Δ_0 in several expressions above is defined as

$$\Delta_0 = \frac{3}{Z} F_0 \,, \tag{11}$$

where $F_0 = 2\pi (4/15)^{2/3} Z_p n_p^{2/3}$ is the Holtsmark normal field strength [16].

In Equation (8), $\phi(\tau)$ is the characteristic function of the probability distribution of the plasma microfield magnitudes $\beta = F/F_0$

$$W(\beta) = \frac{2}{\pi} \beta \int_0^\infty x \sin(\beta x) \exp\left[-\phi(x)\right] dx.$$
 (12)

For a neutral radiator, one can use [17]

I

$$\phi(x) = \frac{x^{3/2}}{(1 + 1.295\sqrt{\Gamma_p x} + 0.606\,\Gamma_p x)},$$
(13)

To summarize, Equations (7)–(13) allow for calculating the shape of the Lyman- $\alpha \pi$ component as broadened by a single plasma species (electrons or ions). The total line shape is obtained by a convolution of the individual contributions [18].

3. Computer Simulations

A variant of computer simulations (CS) described in Ref. [19] is used. Briefly, the Heisenberg equation

$$-i\frac{\partial}{\partial t}\vec{d}(t) = \left[H(t),\vec{d}(t)\right]$$
(14)

is numerically solved by introducing the time-development operator U(t) in the interaction representation,

$$i\frac{d\mathcal{U}(t)}{dt} = \hat{V}_I(t)\mathcal{U}(t)$$
(15)

with

$$\hat{V}_{I}(t) = e^{iH_{0}t} V_{I}(t) e^{-iH_{0}t}$$
(16)

The time evolution of the dipole operator is then given by

$$\vec{d}(t) = \mathcal{U}^{\dagger}(t)e^{iH_0t}\vec{d}e^{-iH_0t}\mathcal{U}(t)$$
(17)

with Fourier transform

$$\vec{d}(\omega) = \int_0^\infty dt e^{i\omega t} \vec{d}(t).$$
(18)

Assuming the radiator density matrix is diagonal, which is customary in line shape broadening calculations [5], the line shape is

$$L(\omega) \propto \sum_{if} \rho_i \left\langle |\vec{d}_{fi}(\omega)|^2 \right\rangle, \tag{19}$$

where the sums are over initial and final states i and f, respectively, and the plasma average denoted by the angle brackets is accomplished by averaging over CS runs.

The motion of the plasma quasiparticles (both plasma electrons and ions) is described by the screened monopole interaction using a velocity Verlet algorithm [20]. However, for the present calculations, the radiators are neutral, therefore, the trajectories of all plasma particles were assumed to be straight.

Each ion species *s* is assigned a different screening length. For a weakly coupled plasma, the inverse screening length κ_s includes screening by all other charged particles with the same or lesser mass,

$$\kappa_s^2 = \sum_{s'(m_{s'} \le m_s)} \frac{4\pi n_{s'} Z_{s'}^2}{kT_{s'}},\tag{20}$$

with m_s , n_s , and T_s the mass, number density, and temperature of species *s* in the plasma.

The simulation follows the reduced-mass model [21] with a fixed, static radiator at the center of a spherical box of radius several times the electron Debye length to ensure convergence [22]. Whenever a perturber exits the simulation volume, it is reinjected at a random point on the sphere surface with a velocity randomly chosen according to the 2D Gaussian distribution in the tangential plane and Rayleigh distribution in the radial direction.

4. Results and Conclusions

In Figure 2, a comparison between the Lyman- $\alpha \pi$ line shapes calculated by the analytical model (Section 2) and computer simulations (Section 3) is shown, indicating a good agreement between the two.



Figure 2. Comparison of the Lyman- $\alpha \pi$ line shape calculated using the computer simulations (CS) and the analytical model. $n_e = 10^{14} \text{ cm}^{-3}$, T = 1 eV, and B = 4 T.

In white dwarf atmospheres, the magnetic field can easily reach hundreds of teslas, e.g., see [23], with the electron density about 10^{17} cm⁻³ to 10^{18} cm⁻³ [24]. In Figure 3, a comparison under such conditions is shown. As in the previous example, a good agreement is seen. In both examples, the values of the full width at half maximum (FWHM) of the lineshapes differ by ~15%.



Figure 3. Same as Figure 2, but for $n_e = 10^{17} \text{ cm}^{-3}$, T = 1 eV, and B = 400 T.

Notably, this value (\sim 15%) of the extraneous width predicted by the model remains almost constant over a very wide range of the plasma densities, as shown in Figure 4. Thus, for practical purposes, it may be suggested to multiply the width given by the model by the 0.85 factor. It is also noted that the model remains sufficiently accurate even beyond its



Figure 4. Comparison of the Lyman- $\alpha \pi$ broadening as given by the computer simulations and the model over wide ranges of densities. *B* = 4 T (**a**) and *B* = 400 T (**b**) are assumed.

To conclude, in this study it is demonstrated that in the strong-*B* limit, when the three components of the Lyman- α Zeeman triplet are well-resolved, the broadening of the central, π component is independent of the magnetic field and, thus, can be used for the plasma density diagnostics, as has recently been suggested [9]. Furthermore, the shape of this component is expressed analytically.

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